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Uncertainty Propagation and Nonlinear Filtering for Space Navigation using Differential Algebra

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-  **Motivations of the Work**
-  **Introduction to Space Navigation**
-  **DA-based higher order Taylor series approach**
 - Nonlinear Uncertainty Propagation
 - Nonlinear Filtering
-  **Simulations and Results**
-  **Conclusions and Development Axes**



- Increase of interplanetary/deep space/sample and return exploration missions



- Long duration
- Large heliocentric/geocentric distances
- Unknown and hostile environments
- Complex dynamics



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➔ Strong need to study the navigation problem in space

A navigation system is a state estimation filter that, starting from sensors measurements, can estimate the spacecraft state variables

- The state of a space vehicle can include a large number of parameters, first of all satellite orbital position and velocity



- The orbit model can be written in the following general equation form:

$$\mathbf{x}_{k+1} = \mathbf{f}(\mathbf{x}_k, \mathbf{t}_k) + \mathbf{w}_k$$

$$\mathbf{y}_k = \mathbf{h}(\mathbf{x}_k, \mathbf{t}_k) + \mathbf{v}_k$$



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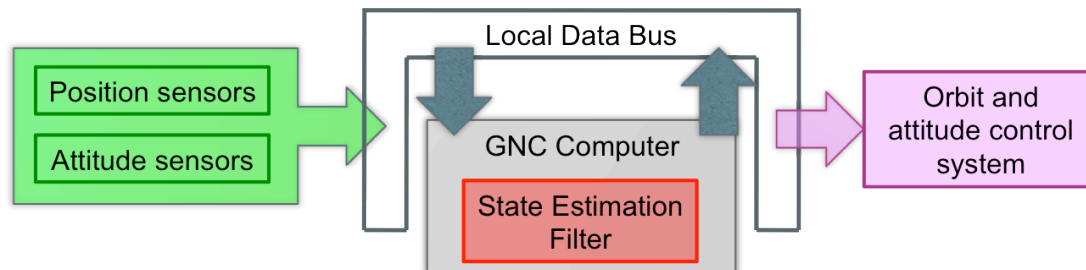
- The on-board navigation system is made up by two main steps:

1. PREDICTION STEP

Use of the dynamics equation to predict the state of the vehicle in a future time

2. CORRECTION STEP

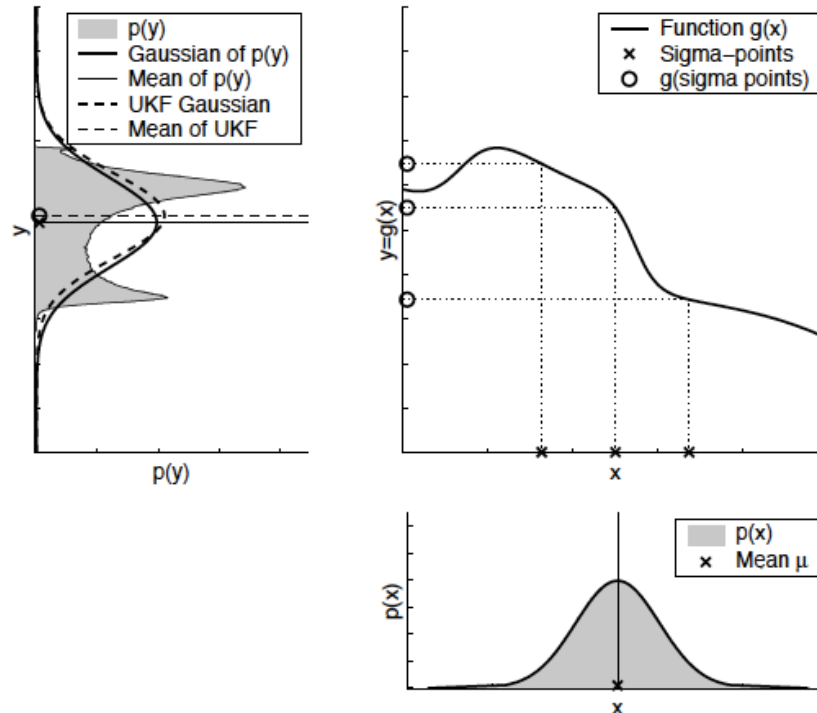
Use the measurements from the on-board sensors (autonomous navigation) or from the ground station (ground-based navigation) to correct the predicted state





- **Autonomous navigation system requirements**

- Real time estimation
- High accuracy
- Need to fully account for nonlinearities in the system





- **Present day orbit uncertainty propagation and filtering**
 - Linearized models
 - Full nonlinear Monte Carlo simulations
 - **Higher order Taylor series approaches**



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- ~~• Linearized models~~ — LOW ACCURACY

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➔ **Differential Algebra can provide powerful tools to face the problem**

- Possibility to **consider nonlinearities**

- Strong **reduction of computational costs**

- Opportunity to **directly evaluate the impact of changes in some physical variables** on the estimated state

- ▶ Impact on the estimated state of errors in the initial state

- ▶ Impact on the estimated state of measurement noises
(sensors architecture definition)



- Given any function f of v variables, DA delivers its Taylor expansion up to the arbitrary order n

- Any integration scheme is based on algebraic operations, involving the evaluation of f at several integration points

➔ Replacing x_0 with $[x_0] = x_0 + \delta x_0$ and carrying out all the operations within the DA framework ➔ Taylor expansion of the ODE flow



$$[x_k] = x_k + \mathcal{M}(\delta x_0, \delta \alpha, \delta \beta, \delta \gamma, \dots)$$

- ➔ Pointwise integration can be replaced by fast evaluation of polynomials
- ➔ The derivatives of the function up to the order of the Taylor expansions are available



Higher order Taylor series approaches



-  **Nonlinear uncertainty propagation**
-  Nonlinear filtering



Nonlinear Mapping of the System Dynamics



- Consider the spacecraft dynamics governed by the **equations of motion**:

$$\begin{cases} \dot{\mathbf{x}}_i = \mathbf{f}_i[t, \mathbf{x}(t)] \\ \mathbf{x}_i(t^0) = \mathbf{x}_i^0 \end{cases}$$



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- Performing a **Taylor series expansion** of the deviation of the current state from the nominal trajectory in terms of the initial deviation:

$$\delta \mathbf{x}_i(\mathbf{t}) = \sum_{p=1}^m \Phi_{i, k_1 \dots k_p} \delta \mathbf{x}_{k_1}^0 \dots \delta \mathbf{x}_{k_p}^0$$



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- The **current mean and covariance matrix** can be written as:

$$\delta \mathbf{m}_i(t) = \sum_{p=1}^m \frac{1}{p!} \Phi_{i,k_1 \dots k_p} \mathbf{E}[\delta \mathbf{x}_{k_1}^0 \dots \delta \mathbf{x}_{k_p}^0]$$

$$\mathbf{P}_{ij}(t) = \sum_{p=1}^m \sum_{q=1}^m \frac{1}{p!q!} \Phi_{i,k_1 \dots k_p} \Phi_{j,\gamma_1 \dots \gamma_q} \mathbf{E}[\delta \mathbf{x}_{k_1}^0 \dots \delta \mathbf{x}_{k_p}^0 \delta \mathbf{x}_{\gamma_1}^0 \dots \delta \mathbf{x}_{\gamma_q}^0]$$



Nonlinear Mapping of the System Dynamics



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$$\left\{ \begin{array}{l} \dot{\Phi}_{i,a} = \mathbf{f}_{i,\alpha}^* \Phi_{\alpha,a} \\ \dot{\Phi}_{i,ab} = \mathbf{f}_{i,\alpha}^* \Phi_{\alpha,ab} + \mathbf{f}_{i,\alpha\beta}^* \Phi_{\alpha,a} \Phi_{\beta,b} \\ \dot{\Phi}_{i,abc} = \mathbf{f}_{i,\alpha}^* \Phi_{\alpha,abc} + \mathbf{f}_{i,\alpha\beta}^* (\Phi_{\alpha,a} \Phi_{\beta,bc} + \Phi_{\alpha,ab} \Phi_{\beta,c} + \Phi_{\alpha,ac} \Phi_{\beta,b}) + \mathbf{f}_{i,\alpha\beta\gamma}^* \Phi_{\alpha,a} \Phi_{\beta,b} \Phi_{\gamma,c} \\ \dots \end{array} \right.$$



Nonlinear Mapping of the System Dynamics



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- The higher order derivatives can be directly extracted from the expansion map
 - ➔ The integration of this system is substituted with ONE integration in DA
 - ➔ More flexible approach
 - ➔ Very general approach (applying nonlinear transformations to mean and covariance estimates)



- The second numerical operation is the **higher-order moment computation**

$$\delta \mathbf{m}_i(t) = \sum_{p=1}^m \frac{1}{p!} \Phi_{i,k_1 \dots k_p} \underline{E[\delta \mathbf{x}_{k_1}^0 \dots \mathbf{x}_{k_p}^0]}$$

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→ Calculated using the **Joint Characteristic Function** of the initial distribution:

$$E[x_i] = m_i$$

$$E[x_i x_j] = m_i m_j + P_{ij}$$

$$E[x_i x_j x_k] = m_i m_j m_k + (m_i P_{jk} + m_j P_{ik} + m_k P_{ij})$$

$$E[x_i x_j x_k x_l] = m_i m_j m_k m_l + (m_i m_j P_{kl} + m_i m_k P_{jl} + m_j m_k P_{il} + m_i m_l P_{jk} + m_j m_l P_{ik} + m_k m_l P_{ij}) + P_{ij} P_{kl} + P_{ik} P_{jl} + P_{il} P_{jk}$$

NOTE: If the initial mean is zero → Further simplification: **all the odd moment are zero**



Nonlinear Mapping of the System Dynamics

- Consider the two-body problem of a Earth graviting satellite

- Initial state defined as DA variable

(lengths units scaled by the semimajor axis and time units by the factor $\sqrt{\frac{a^3}{\mu}}$)

$$\mathbf{x}_0 = \begin{pmatrix} \mathbf{r}_0 \\ \mathbf{v}_0 \end{pmatrix} = \begin{pmatrix} -0.68787 + \delta x \\ -0.39713 + \delta y \\ +0.28448 + \delta z \\ -0.51331 + \delta v_x \\ +0.98266 + \delta v_y \\ +0.37611 + \delta v_z \end{pmatrix} \quad \mathbf{P} = \begin{bmatrix} 0.01 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0.01 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0.01 & 0 & 0 & 0 \\ 0 & 0 & 0 & 10^{-4} & 0 & 0 \\ 0 & 0 & 0 & 0 & 10^{-4} & 0 \\ 0 & 0 & 0 & 0 & 0 & 10^{-4} \end{bmatrix}$$

- Carry out the integration of the motion in differential algebra
- Calculate the estimation of the mean and covariance



Nonlinear Mapping of the System Dynamics

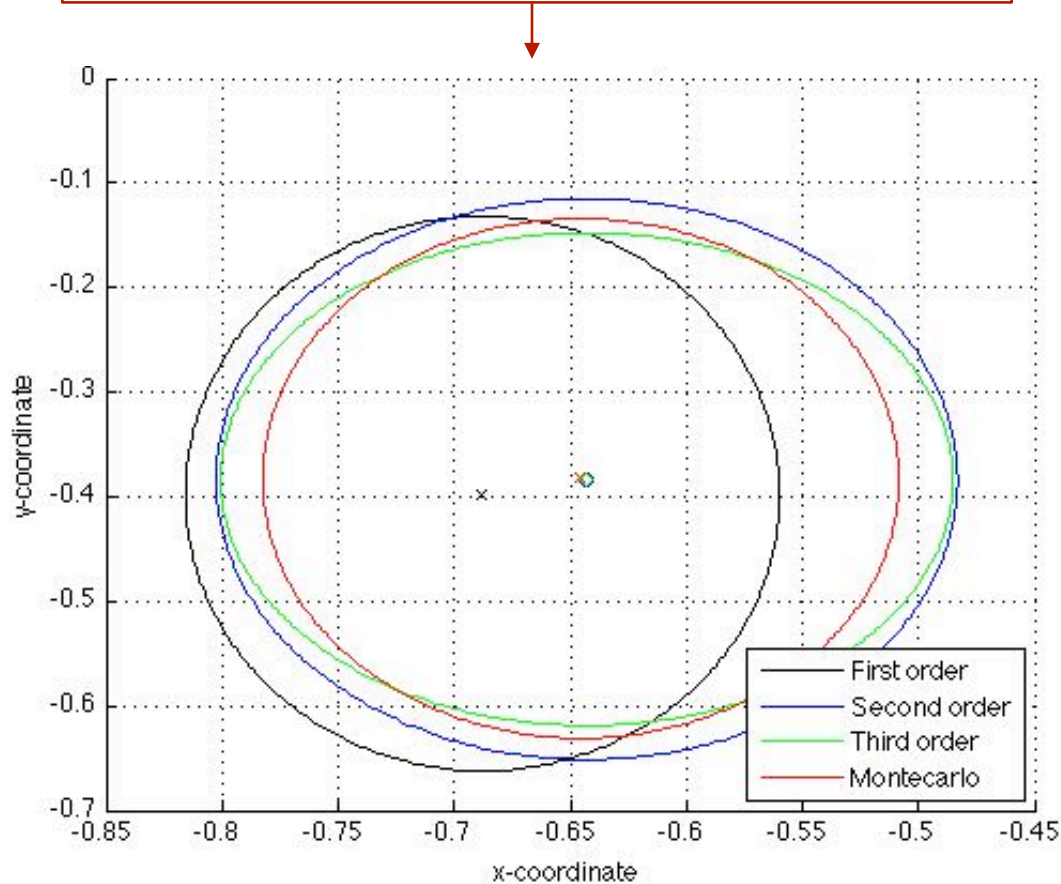


	DA-1st order	DA-2nd order	DA-3rd order
	$\Delta\mathbf{r}[\%]$	$\Delta\mathbf{r}[\%]$	$\Delta\mathbf{r}[\%]$
0.1 orbit	0.021	0.017	0.017
0.5 orbit	0.559	0.111	0.111
0.8 orbit	3.299	0.007	0.007
1 orbit	6.197	0.092	0.092



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Higher order Taylor series approaches



 Nonlinear uncertainty propagation

 **Nonlinear filtering**



Higher Order Taylor Series Filtering Methods



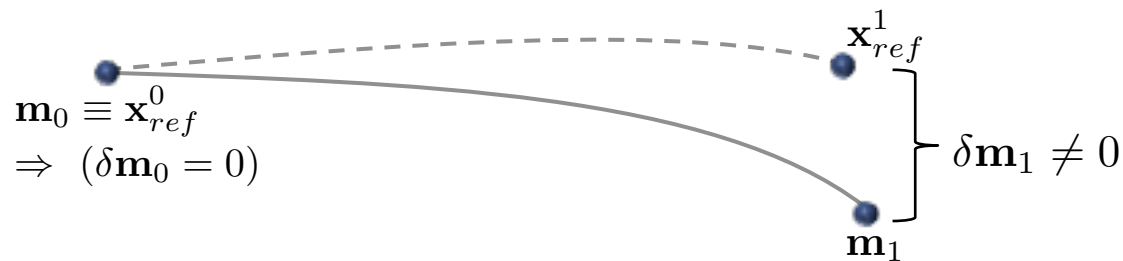
- The application of higher order Taylor series methods to nonlinear filtering is the Higher-Order Numerical Extended Kalman Filter (HNEKF) [*Park&Scheeres2007*]



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HNEKF PREDICTION EQUATION

$$\bullet (m_{k+1}^-)^i = E[\Phi^i(t_{k+1}; \mathbf{m}_k^+ + \delta \mathbf{x}_k, t_k) + w_k^i] = \Phi^i(t_{k+1}; \mathbf{m}_k^+, t_k) + \sum_{p=1}^m \frac{1}{p!} \Phi_{(t_{k+1}, t_k)}^{i, \gamma_1 \dots \gamma_p} E[\delta x_k^{\gamma_1} \dots \delta x_k^{\gamma_p}]$$

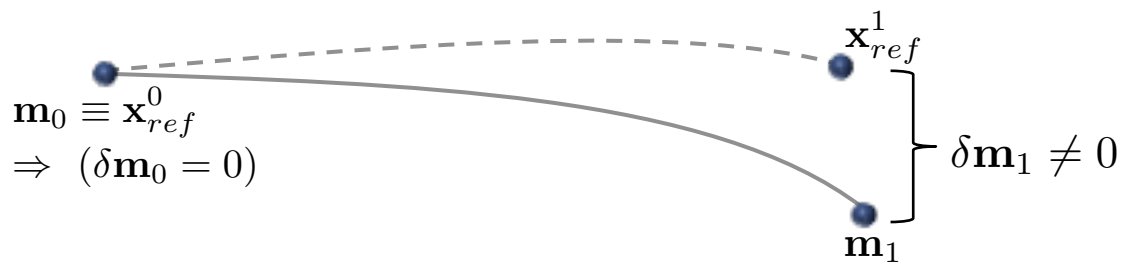




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$$\bullet (P_{k+1}^-)^{ij} = \left(\sum_{p=1}^m \sum_{q=1}^m \frac{1}{p!q!} \Phi_{(t_{k+1}, t_k)}^{i, \gamma_1 \dots \gamma_p} \Phi_{(t_{k+1}, t_k)}^{j, \zeta_1 \dots \zeta_q} E[\delta x_k^{\gamma_1} \dots x_k^{\gamma_p} x_k^{\zeta_1} \dots x_k^{\zeta_q}] \right) - \delta m_{k+1}^i (\delta \mathbf{x}_k) \delta m_{k+1}^j (\delta \mathbf{x}_k) + Q_k^{ij}$$

$$\bullet (n_{k+1}^-)^i = E[h^i(t_{k+1}; \mathbf{m}_k^+ + \delta \mathbf{x}_k, t_k) + v_{k+1}^i] = h^i(t_{k+1}; \mathbf{m}_k^+, t_k) + \sum_{p=1}^m \frac{1}{p!} h_{t_{k+1}, t_k}^{i, \gamma_1 \dots \gamma_p} E[\delta x_k^{\gamma_1} \dots \delta x_k^{\gamma_p}]$$



HNEKF UPDATE EQUATION

- $(P_{k+1}^{zz})^{ij} = R_{k+1}^{ij} + \sum_{p=1}^m \sum_{q=1}^m \frac{1}{p!q!} h_{t_{k+1}, t_k}^{i, \gamma_1 \dots \gamma_p} h_{t_{k+1}, t_k}^{i, \zeta_1 \dots \zeta_q} E[\delta x_k^{\gamma_1} \dots \delta x_k^{\gamma_p} \delta x_k^{\zeta_1} \dots \delta x_k^{\zeta_q}] - (\delta n_{k+1}^-)^i (\delta n_{k+1}^-)^j$
- $(P_{k+1}^{xz})^{ij} = \sum_{p=1}^m \sum_{q=1}^m \frac{1}{p!q!} \Phi_{t_{k+1}, t_k}^{i, \gamma_1 \dots \gamma_p} h_{t_{k+1}, t_k}^{i, \zeta_1 \dots \zeta_q} E[\delta x_k^{\gamma_1} \dots \delta x_k^{\gamma_p} \delta x_k^{\zeta_1} \dots \delta x_k^{\zeta_q}] - (\delta m_{k+1}^-)^i (\delta n_{k+1}^-)^j$
- $\mathbf{K}_{k+1} = \mathbf{P}_{k+1}^{xz} (\mathbf{P}_{k+1}^{zz})^{-1}$
- $\mathbf{m}_{k+1}^+ = \mathbf{m}_{k+1}^- + \mathbf{K}_{k+1} (\mathbf{z}_{k+1} - \mathbf{n}_{k+1}^-)$
- $\mathbf{P}_{k+1}^+ = \mathbf{P}_{k+1}^- - \mathbf{K}_{k+1} \mathbf{P}_{k+1}^{zz} \mathbf{K}_{k+1}^T$



HNEKF UPDATE EQUATION

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- The **state** and the **measurements** have been defined as **DA variables**

$$[\mathbf{x}_{k+1}] = \bar{\mathbf{x}}_{k+1} + \mathcal{M}(\delta \mathbf{x}_k)$$

$$[\mathbf{z}_{k+1}] = \bar{\mathbf{z}}_{k+1} + \mathcal{M}(\delta \mathbf{x}_{k+1}) = \bar{\mathbf{z}}_{k+1} + \mathcal{M}(\delta \mathbf{x}_k)$$



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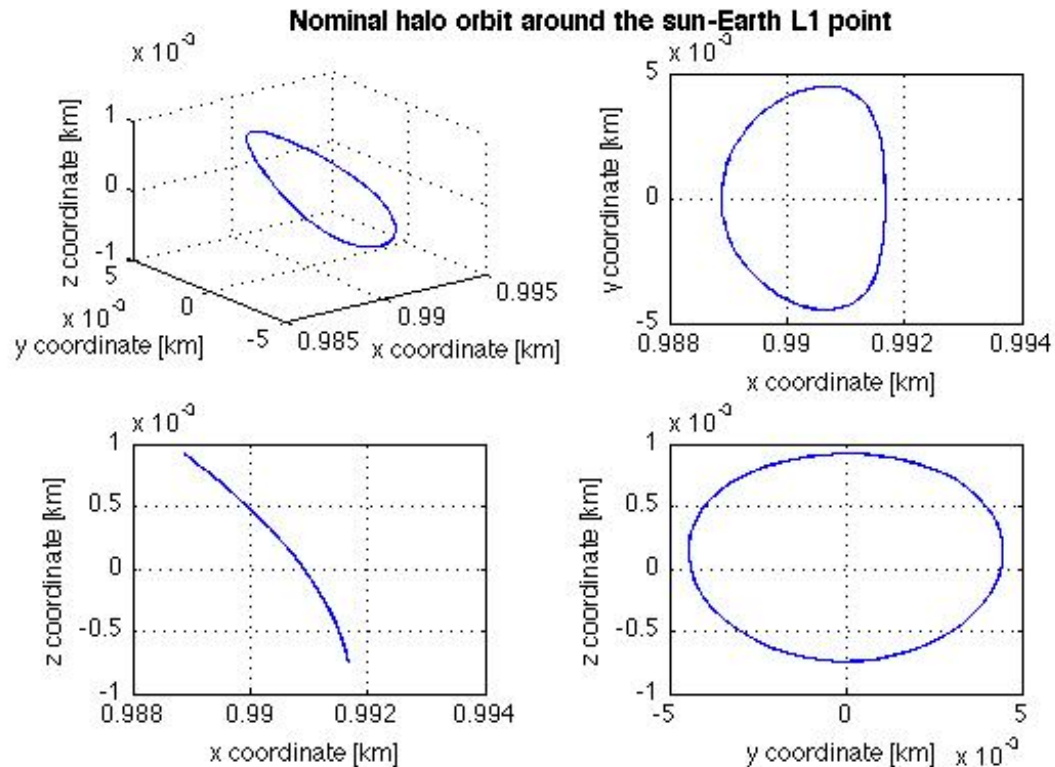
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- Higher order derivatives are **provided from one DA integration of the state and measurement computation**



HNEKF: Halo orbit about the Sun-Earth L_1 point

- **State:** position and velocity
- **Measurement model:** linear
- **Initial guess:** off by 100 km and 0.1 m/s from the true state (boundary of initial ellipsoid)
- **No process noise:** errors in the initial state and in the measurements

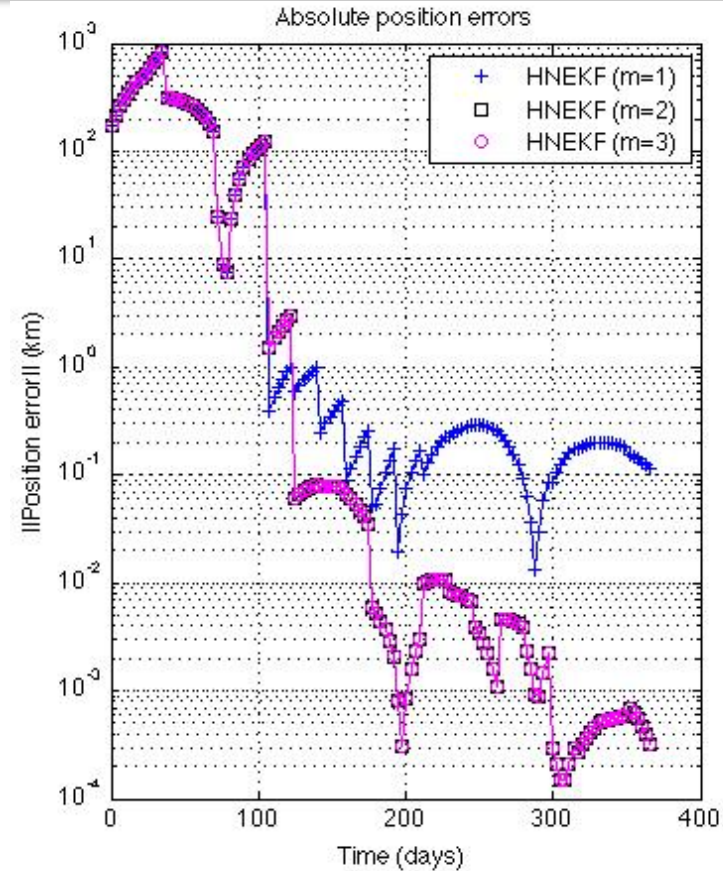
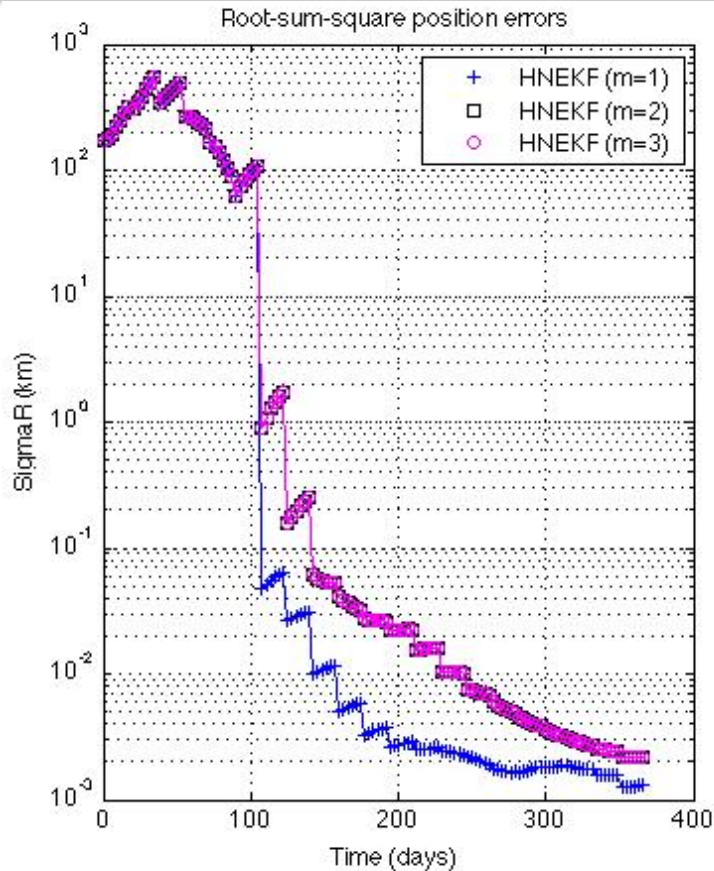




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- **Without using DA:**

- Two system of differential equations up to the m^{th} (order of the Taylor expansion) order must be solved
- If something changes in the dynamical or measurement equations the whole solving system must be rewritten

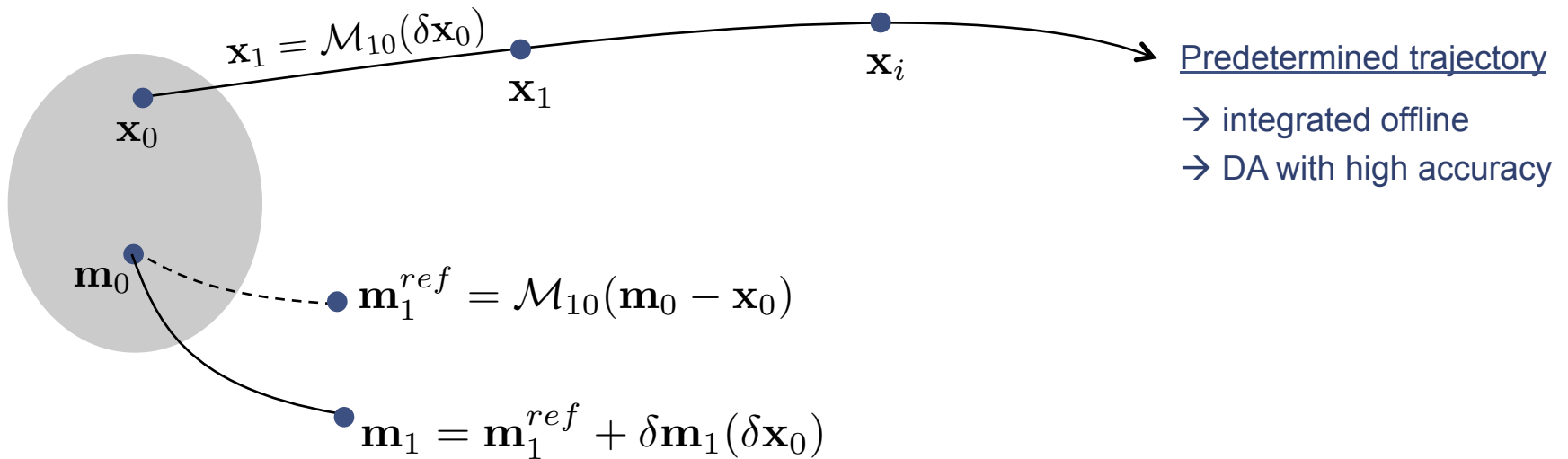
- **Using DA:**

- At each time step only one integration in DA is required
- More flexibility
- The same accuracy of the standard HNEKF is guaranteed



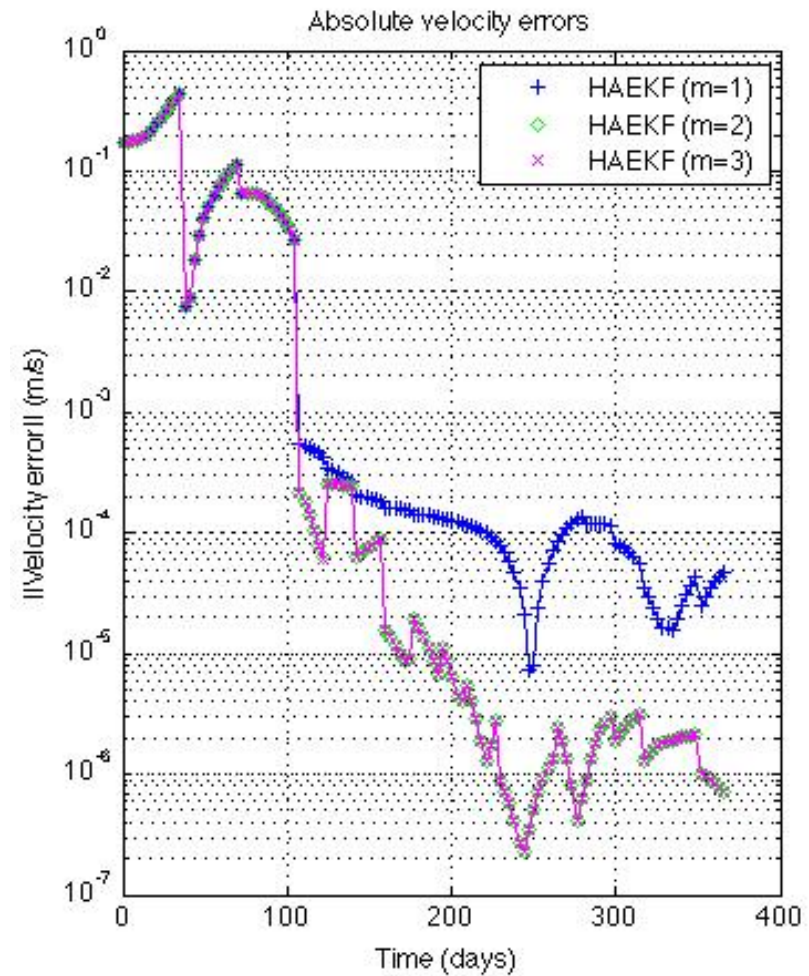
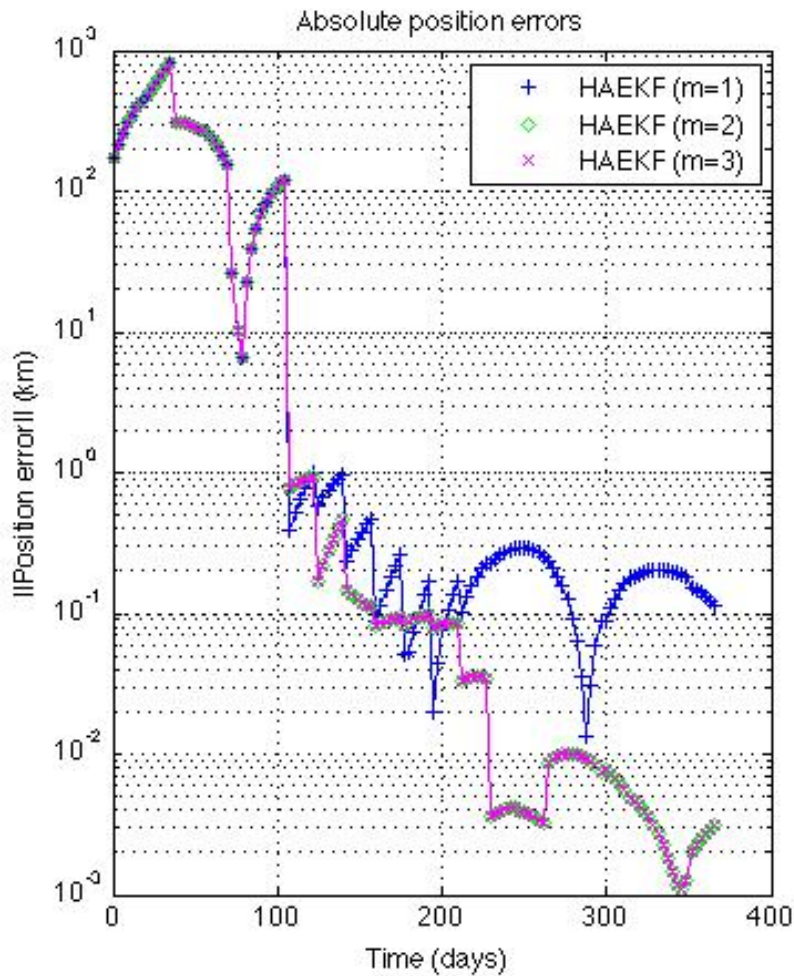
- Missions with predetermined reference trajectories

→ compute the reference trajectory over a time span before filtering



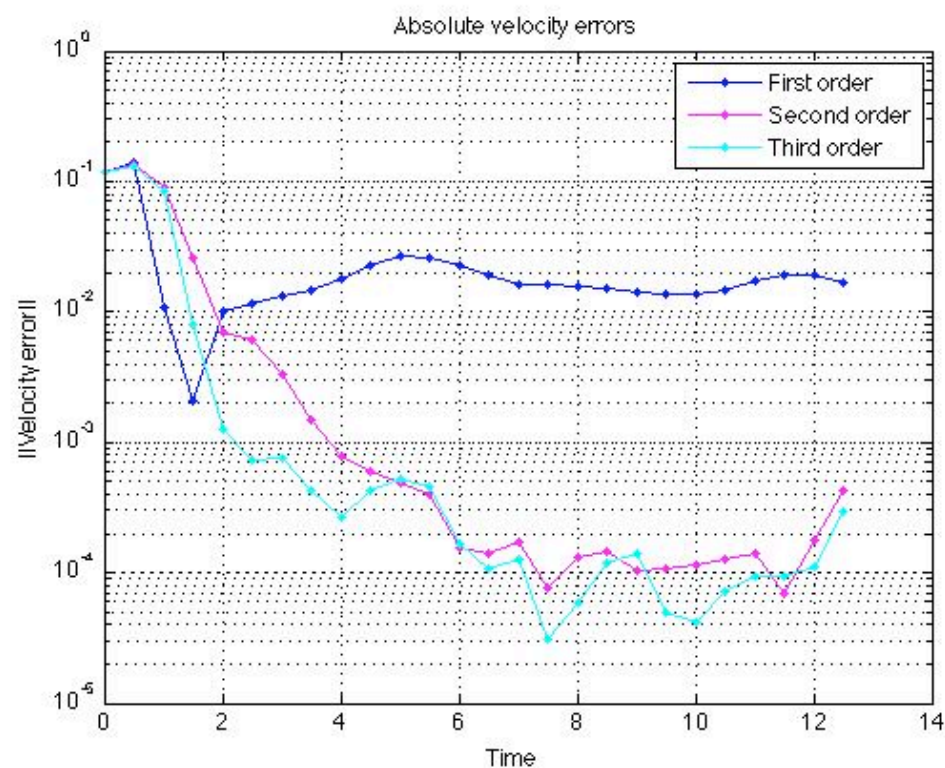
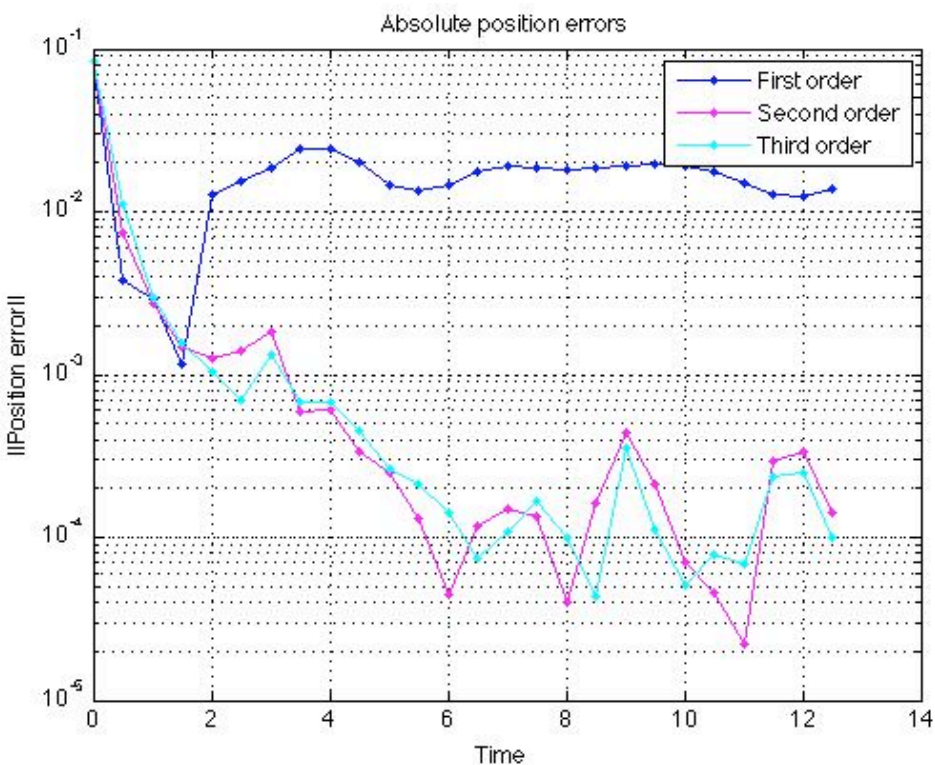


HAEKF: Halo orbit about the Sun-Earth L_1 point





- DA-based HNEKF behavior in case of low frequency estimation and nonlinear measurements
 - Measuring the velocity from S/C to Earth, the right ascension and declination of the Earth
 - Better performance of higher orders wrt first order filter





CONCLUSIONS

- New methods for nonlinear uncertainty propagation and filtering have been presented
- The methods are based on Taylor differential algebra implemented in COSY-Infinity
- The methods have good performance in terms of computational costs and flexibility

FURTHER DEVELOPMENTS

- Study and development of Monte Carlo-based filters using DA
- Use of DA to define a sensors architecture that increase estimation precision and reduce the computational load



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Uncertainty Propagation and Nonlinear Filtering for Space Navigation using Differential Algebra

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